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| 1. Record Nr.           | UNINA9911019703903321  |
| Autore                  | Rudiger G (Gunther)  |
| Titolo                  | The magnetic universe : geophysical and astrophysical dynamo theory /<br>/ Gunther Rudiger and Rainer Hollerbach   |
| Pubbl/distr/stampa      | Weinheim, : Wiley-VCH, c2004   |
| ISBN                    | 9786610520152<br>9781280520150<br>1280520159<br>9783527603657<br>3527603654<br>9783527605002<br>3527605002   |
| Descrizione fisica      | 1 online resource (346 p.)   |
| Altri autori (Persone)  | HollerbachRainer   |
| Disciplina              | 523.01/886   |
| Soggetti                | Dynamo theory (Cosmic physics)<br>Cosmic magnetic fields<br>Magnetohydrodynamics<br>Astrophysics   |
| Lingua di pubblicazione | Inglese  |
| Formato                 | Materiale a stampa   |
| Livello bibliografico   | Monografia   |
| Note generali           | Description based upon print version of record.  |
| Nota di bibliografia    | Includes bibliographical references and index.   |
| Nota di contenuto       | The Magnetic Universe Geophysical and Astrophysical Dynamo Theory;<br>Contents; Preface; 1 Introduction; 2 Earth and Planets; 2.1<br>Observational Overview; 2.1.1 Reversals; 2.1.2 Other Time-Variability;<br>2.2 Basic Equations and Parameters; 2.2.1 Anelastic and Boussinesq<br>Equations; 2.2.2 Nondimensionalization; 2.3 Magnetoconvection; 2.3.1<br>Rotation or Magnetism Alone; 2.3.2 Rotation and Magnetism Together;<br>2.3.3 Weak versus Strong Fields; 2.3.4 Oscillatory Convection Modes;<br>2.4 Taylor's Constraint; 2.4.1 Taylor's Original Analysis; 2.4.2<br>Relaxation of $Ro = E = 0$<br>2.4.3 Taylor States versus Ekman States2.4.4 From Ekman States to<br>Taylor States; 2.4.5 Torsional Oscillations; 2.4.6 -Dynamoes; 2.4.7<br>Taylor's Constraint in the Anelastic Approximation; 2.5 Hydromagnetic<br>Waves; 2.6 The Inner Core; 2.6.1 Stewartson Layers on C; 2.6.2<br>Nonaxisymmetric Shear Layers on C; 2.6.3 Finite Conductivity of the |

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## Sommario/riassunto

Magnetism is one of the most pervasive features of the Universe, with planets, stars and entire galaxies all having associated magnetic fields. All of these fields are generated by the motion of electrically conducting fluids, the so-called dynamo effect. The precise details of what drives the motion, and indeed what the fluid consists of, differ widely though. In this work the authors draw upon their expertise in geophysical and astrophysical MHD to explore some of these phenomena, and describe the similarities and differences between different magnetized objects. They also explain why magn