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Nota di contenuto	Series Editors; Advisory Board; Preface; Contents; Contributors; 1 Spring School of Spectroscopic Data Analyses: Determination of Atmospheric Parameters of B, A, F and G-type Stars---Introduction; 1 The Need for Accurate Atmospheric Parameters and Abundances from an Astrophysical Point of View; 2 Spectra as a Powerful Tool; 2.1 How to Choose the Ideal Spectrograph?; 3 KASC and the Characterisation of Asteroseismic Kepler Targets; 4 Motivation of the Spectroscopic School; 4.1 Available Spectra of Kepler Main Sequence Pulsators; 4.2 Outlook; References 2 Stellar Atmospheres: Basic Processes and Equations1 Introduction; 2 Basic Definitions; 2.1 Specific Intensity; 2.2 Flux; 2.3 K-integral; 3 Absorption Coefficient and Optical Depth; 4 Emission Coefficient; 5 Source Function and Its Physical Meaning; 5.1 Two Simple Cases; 5.2 Pure Isotropic Scattering; 5.3 Pure Absorption; 6 The Transfer Equation; 6.1 Elementary Solutions; 6.2 No Absorption, No Emission; 6.3 No Absorption, Only Emission; 6.4 No Emission, Only Absorption; 6.5

General Case: Absorption and Emission; 6.6 Special Case: Linear Source Function; 7 The Flux Integral

8 Computing a Model Atmosphere

8.1 Basic Equations of Stellar Atmospheres; 8.2 A Pedagogical Example: The Grey Atmosphere; 9 Conclusions; References; 3 How to Build a Model of the Atmosphere and Spectrum; 1 Introduction; 2 Observational and Computational Pipelines; 3 Rotation; 4 Circumstellar and Interstellar Absorption; 5 Telluric Spectra; 6 Comparison to Observed Spectra; 7 Equation of State; 8 Convection and Microturbulent Velocity; References; 4 Model Atmosphere Codes: ATLAS12 and ATLAS9; 1 Introduction; 2 SYNTHE; 3 High Resolution Atlases; 4 Validating Spectrum Calculations; 5 ATLAS12

6 DFSYNTHE and ODFs

7 ODFs and ATLAS9; References; 5 Atomic Data: Where to Get Them, How to Use Them; 1 Introduction; 2 Spectral Line Formation; 3 Atomic Line Databases; 3.1 Robert Kurucz Database; 3.2 NIST Atomic Spectra Database; 3.3 VALD3---Vienna Atomic Line Database, Third Release; 3.4 D.E.S.I.R.E. and D.R.E.A.M. Databases at Mons University; 3.5 Stark-b Database; 4 VAMDC---Virtual Atomic and Molecular Data Centre; References; 6 Problems with Atomic and Molecular Data: Including All the Lines; 1 Introduction; 2 Problems; 3 Examples of New Calculations; 4 Molecules, TiO and H<sub>2</sub>O

5 Spectrum Analysis Using Stellar Atlases as the Laboratory Source

6 Conclusion; References; 7 Spectral Classification: The First Step in Quantitative Spectral Analysis; 1 Introduction; 2 The Role of Classification in Spectral Analysis; 3 How are Spectral Types Determined?; 3.1 Standard Stars; 3.2 An Example of How to Classify a Star; 4 Conclusions; References; 8 Effective Temperature Determination; 1 Effective Temperature; 2 The Paschen Continuum; 3 Temperatures from Balmer Line Profiles; 3.1 Examples of Balmer Line Fitting; 4 Spectral Line Depth Ratios; 5 Equivalent Width

6 Metal Line Diagnostics

## Sommario/riassunto

This book introduces the theory of stellar atmospheres. Almost everything we know about stars is by analysis of the radiation coming from their atmospheres. Several aspects of astrophysics require accurate atmospheric parameters and abundances. Spectroscopy is one of the most powerful tools at an astronomer's disposal, allowing the determination of the fundamental parameters of stars: surface temperature, gravity, chemical composition, magnetic field, rotation and turbulence. These can be supplemented by distance measurements or pulsation parameters providing information about stellar interior and stellar evolution, otherwise unavailable. The volume is based on lectures presented at the Wroclaw's Spectroscopic School aimed at training young researchers in performing quantitative spectral analysis of low-, mid-, and high-resolution spectra of B, A, and F-type stars.